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FROM THE PUBLISHER

For subscription purposes, this is the third issue of 1981.

Please note the new subscription and membership rates spelled out in the paragraph at the top of the second column, on this page. Air mail shipment of o.w. subscriptions remains at \$1.80/year extra, outside the U.S.A., Canada, and Mexico. Back issues still are priced at \$1.00 per issue, but this probably will be changed in the near future.

European and U. K. observers should join IOTA/ES, sending DM 12.-- to Hans J. Bode, Bartold-Knaust Str. 8, 3000 Hannover 91, German Federal Republic.

IOTA NEWS

David W. Dunham

I am starting to write this only a few hours before departing for Somalia to observe the occultation of Nunki by Venus on November 17, so my written contributions for this issue necessarily are brief. We will need to produce the next issue in December; it will include details about the total eclipse of the moon on January 9, 1982, visible from the Eastern Hemisphere. Of the three total lunar eclipses of 1982, the richest star field is traversed on January 9. The 3.5-magnitude star δ Geminorum (Wasat = Z.C. 1110) will be occulted during totality in southeastern Asia and northwestern Australia; it is the brightest star to be occulted during a total lunar eclipse until 1985. The northern limit lies about 100 km northeast of Manila, Philippines, while the southern limit crosses Australian desert areas. It could be the first occultation wherein grazes are observed at both the north and south limbs, which would be valuable for accurately determining the moon's polar radius. I plan to distribute detailed predictions of occultations of Astrographic Catalog stars down to about 12th magnitude during the eclipse; especially if you have not received such predictions (J or M catalog) from me before and are in the Eastern Hemisphere, you should request these predictions and send me accurate geographical coordinates of your observing site. My address has changed, but only because our post office has moved to a new building. It is now: P. O. Box 7488, Silver Spring, MD 20907, U.S.A. This is the same as before, except that 7000 has been added to the box number. Reports of grazes and asteroidal occultations or close approaches should be sent to the new address.

John Phelps reports that inflation finally has caught up with us. Increased costs, especially those caused by the two postal rate increases this year and the fact that o.n. now is amounting to about 60 pages per year rather than 40, makes it necessary for us to raise our dues substantially, or we will be in trouble with our finances by the middle of next year. Effective after the publica-of this issue (Number 13), IOTA annual dues will be increased to \$11.00 for North Americans, and \$16.00 for others. Separate subscriptions to o.n. will cost \$5.50, with an additional amount for overseas airmail. We are sorry that these increases are necessary, but we can not keep going at the original 1975 IOTA rates.

Those who have expressed an opinion have been favorably disposed to the idea of incorporation, so we are starting the procedure to incorporate in the State of Illinois. As soon as that is accomplished, we will apply for tax-exempt status as an educational organization.

Besides this article, my only other one for this issue will be the start of the article on predictions of planetary and asteroidal occultations during 1982. Observational details of asteroidal and lunar grazing occultations, and double stars, will be delayed to issue number 15, early in 1982. I will mention briefly that three asteroidal occultations have been observed since the last issue. though in each case, coverage was so poor that a reliable diameter determination will not be possible. Unfortunately, most of the paths shifted away from populous parts of North America, according to astrometry mainly by Klemola at Lick Observatory. The occultation by (18) Melpomene on August 7 was observed by David Herald at Mt. Stromlo Observatory in Australia, and at two telescopes on Mauna Kea, Hawaii. The depths of these photoelectric records were not as much as expected, indicating that perhaps only one component of a close double was occulted. A visual observer on Oahu, about 400 km from the path, timed an 18-second occultation using a 13-inch reflector. Although the magnitude drop was small, the altitude was high and the seeing very good. The observation indicates a separation of the stellar components of about 1.2, for SAO 145972. On October 5, an occultation by (105) Artemis was recorded at Sutherland, South Africa, after being telephoned by Gordon Taylor, who just recently had obtained a plate of the approaching objects, only a few hours before.

There was about a 0.75 southward shift for the oc-

cultation by (88) Thisbe on October 7, which was predicted to about one radius accuracy from measurements of a plate taken at the U.S. Naval Observatory at Flagstaff, AZ, the night before. About a dozen observations were made of the occultation, mainly by observers in the Denver and Minneapolis areas close to the center of the actual path; the longest durations were 11 seconds, about 10% greater than expected. Paul Maley and James Fox traveled well north of the others to obtain valuable northern chords, but no data were obtained on the southern half of Thisbe, which leaves its shape poorly determined, rather similar to the situation for the occultation by (78) Diana on 1980 September 4. An interesting trailed photograph obtained at the Barnard Observatory in Golden, Colorado, shows the occultation; it probably will be published in the January issue of Sky and Telescope.

Alan Fiala and Robert Bolster recorded ten events during the grazing occultation of 6.8-mag. ZC 796 at Stormont, VA, on September 20. They used the same video camera that Alan used during the graze of δ Cancri last May (p. 167 of the last issue) and the 14-inch Schmidt-Cassegrain owned by the National Capital Astronomers. The star was near the limits of detectibility with the video system; six other visual observers timed the graze. Mark Trueblood used a similar camera, which I recently purchased, with his video recorder and 8-inch Schmidt-Cass at a location about a kilometer farther north, where a miss occurred. The star was only faintly visible on the monitor, and almost impossible to record. Video suffers from the same contrast problems that plague visual observers during gibbous-phase grazes, and stars of similar brightness to ZC 796 probably could be recorded easily when the moon is a thin crescent. A week later, Mark Trueblood and I set up the same equipment on a moonless night. We got a reasonably good signal from 7th-magnitude stars, and could just barely detect 8th-magnitude stars. The R.C.A. TC 2055 Ultracon costs less than the price we mentioned in the last issue. Prices vary, with the lowest around \$700 for the camera. Most video stores do not have it in stock, and have to wait 30 to 45 days to get one from the factory in Lancaster, PA. A good portable video recorder-player is in the same price range.

Studies of the orbital elements of some of the asteroids involved in 1982 occultations (see article starting next column) have shown that the Herget data for (386) Siegena are much better than those in the E.M.P.'s. Consequently, for the occultation on 1981 Nov. 22, the primary's path will miss the earth's surface, while the occultation of the 9.7-magnitude secondary probably will cross the U.K. and parts of western Europe. Unfortunately, both Gordon Taylor and I still will be in Africa for the occultation of Nunki by Venus, so that we will not be available to compute an astrometric improvement of the prediction.

A CORRECTION TO THE 1977 TOTAL OCCULTATION TALLY

Joseph E. Carroll

The 1977 occultation tally (o.N. 2 (12), 156 ff.) contained an error which deprived N. P. Wieth-Knudsen of the number one position for individual observers. He is, as I have found out, a single observer and he has always been. Therefore, the "ET

AL" which appears after his name should be deleted in the referenced tally.

Dr. Wieth-Knudsen is thus the leading lone observer for 1977, and, in fact, has also been among the top ten (tops of pp. 158-9) observers for the last several years (and continues to be thru 1980 according to the data on hand so far).

Also, I have spelled his name wrong. I thus very much apologize for these errors and hope that this correction will be noted by all interested parties.

PLANETARY OCCULTATION PREDICTIONS FOR 1982

David W. Dunham

Predictions of occultations of stars by major and minor planets during 1982 are given in two tables. They are like the tables for the 1981 events; see $o.n.\ 2$ (10) 115-118.

Most of the events in the table were found by Gordon Taylor at the Royal Greenwich Observatory and published in his Bulletin 26 of the I.A.U. Commission 20's Working Group on Predictions of Occultations by Satellites and Minor Planets. Derek Wallentinsen, comparing the SAO catalog manually with astrometric ephemerides supplied by me, independently found many of these events and published predictions of them earlier in Contribution No. 2 of the James-Mims Observatory; the events on the following 1982 dates were not in Taylor's list: Mar. 9, Mar. 19, June 5, and Oct. 27. The events involving Pallas on Mar. 21, Juno on June 11, and (10) Hygiea on Dec. 6 were not in Taylor's Bulletin 26, but were listed in his article about occultations by the four largest minor planets during this decade in Astron. J. 86, p. 903. Drs. L. Wasserman, E. Bowell, and R. Millis at Lowell Observatory have made computer comparisons of ephemerides of 89 asteroids, mostly those larger than 150 km in diameter, with the AGK3 and SAO catalogs to find the following events not in Taylor's lists, or Wallentinsen's: Mar. 3, Mar. 20 (Bettina), April 29, June 4, June 29, and July 14 (both events). The Lowell astronomers have submitted their results for both 1982 and 1983 to the Astronomical Journal, which plans to publish them in the 1981 December issue. They asked Brian Marsden to check the orbital elements published in the Ephemerides of Minor Planets for 1980 (E.M.P. 80) by comparison with recent observations, for the asteroids they were considering. About ten of the orbits were quite poor, so Marsden requested updated elements from Paul Herget in Cincinnati. Herget calculated them, Marsden found them to be much better than the elements in the E.M.P.'s, and the Lowell astronomers used them. Most of the Lowell events not found by Taylor involved the newlyimproved orbits. Unfortunately, only a month after this work was done, Paul Herget passed away.

As was the case for the 1981 events, about half of the asteroidal ephemerides I use are computed from osculating orbital elements computed by Herget and published in the Minor Planet Circulars (M.P.C.'s). Most of the others were calculated at the Leningrad Institute of Theoretical Astronomy (I.T.A.) and published in the E.M.P.'s for 1980 or 1981. For many events, elements from both sources are available, so that I can make two separate predictions. For events on the following dates, the ephemerides dif-

fer by less than 0.4 and less than 2.5 minutes in time, smaller than the expected errors relative to the occulted stars: Jan. 21, Jan. 24, Mar. 24, Apr. 27, July 4, Aug. 29, Sept. 15, Sept. 21 (324), Oct. 7, Oct. 8, Nov. 18, and Dec. 13. Larger ephemeris differences are given in the final table, in the

sense I.T.A. minus Herget, except for the noted cases. The value in the shift column gives the path differences in arc seconds measured perpendicular to the asteroid's geocentric motion; the letter following it tells which direction the occultation path will be displaced on the earth's surface from the

nominal (usually Herget) prediction. The value in the Δt column tells whether the geocentric time of closest approach will be early (negative) or late (positive) in minutes relative to the nominal prediction. Some of the differences were so large that I compared the ephemerides with late 1970's (and some early 1980's) observations published in recent numbers of the M.P.C.'s, in particular, for minor

planets 56, 62, 117, 145, 148, 334, 375, 386, 602, and 702. In addition, the I.T.A. orbits for (409) Aspasia and (476) Hedwig were found to be deficient a year ago for 1981 occultations, and Taylor predicts that these objects will occult stars again in 1982. For (409) Aspasia, the observations clearly favor Herget's orbit over the I.T.A. orbit, which was used by Taylor for his consequently incorrect

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Herget's path misses the earth by 0.4 to the south, but since even his orbit is probably accurate to only ± 2 ", an occultation is possible in South Africa. All available orbits for (476) Hedwig are very poor, and have not been improved since last year, so it is none none none w145 W w105 W w 80 W a11 a11 w115 E w120 M e 40 E 0 (%Sn1 $\Sigma \square$ 112° 117° 156 156 1172 1172 1173 1170 1170 1170 1170 1170 1170 3 Indonesia, Tanzania
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prediction of the occultation on 1982 August 1.

on April 8 will not be visible from the earth's surface. Also, the observations clearly favor Herget's orbit over the I.T.A. orbits for (216) Kleopatra, (334) Chicago, (386) Siegena, (602) Marianna, and (702) Alauda, so that Taylor's I.T.A.-based predictions for these objects are wrong. In fact, Taylor's occultations predicted by Kleopatra on May 28, Alau-z A ۵, R.A. АР COMPARISON DATA
K3 No Shift Time -0.29 -0.95 -0.82 -0.51 -0.18 $0.64 \\ 0.87$ -0.83 -0.83 -0.61 0 8 <u>8</u> SES S 7.0 0.7 0.3 0.6 0.6 ďf DIAMETER 36 7 7 7 7 19 19 23 37 37 251 416 133 329 401 168 3261 72 163 591 68 175 0.08 0.17 0.06 0.02 0.33 0.13 0.27 0.09 0.09 0.09 0.09 0.09 0 \simeq 용 ⋖ ¥. 88537 29289 98482 55791 55766 98369 48612 87719 38236 39019 운 147137 6017 93544 10157 6017 SAO "/Day PA 60 261 102 258 239 100 253 251 83 261 261 261 0.213 0.110 0.203 0.230 0.330 0.224 0.222 0.223 0.223 0.202 0.202 0.203 0.003 Type 1041 568 617 268 1327 339 1015 955 426 481 800 428 1123 1123 1760 864 E T RSOI 256 0.26 149 0.11 92 0.05 92 0.05 92 0.05 92 0.05 92 0.05 92 0.05 92 0.09 92 0.00 92 0 A . R P L km-diam. Io Parthenope 19 Gallia 24 Egeria 24 Emita 16 Bamberga 2 Undina 1 Wratislavi Lutetia Ursula Ursula Themis Minerva Germania Aquitania 0 Hygiea Irene Emita I N (

quite probable that Taylor's predicted event for it

da on March 28, and Marianna on April 6 will not occur. Also, since Taylor issued his predictions, new orbits were published in E.M.P. 81 by I.T.A. for (62) Erato, (117) Lomia, and (148) Gallia; these new orbits clearly were favored by recent observations. Consequently, Taylor's events involving Lomia on March 16 and Erato on December 20 also will not occur from earth. On the other hand, the observations favored the I.T.A. orbits over Herget's for (56) Melete, (145) Adeona, (344) Desiderata, and (375) Ursula. Since I originally had selected Herget's elements for my predictions involving occultations by these asteroids, I had to change my nominal calculations to the I.T.A. bases, after which my paths were in good agreement with Taylor's. These events are indicated with asterisks in the table below.

Table of Large Ephemeris Differences for 1982

Date	MP#	Shift_	Δt	Notes							
Jan 11* Feb 8*	344 344	2.25 S 2.03 S	+5.8 +4.6	Herget MPC 4368-ITA 1977 Herget MPC 4371-EMP 81 Herget MPC 4371-EMP 81							
Mar 23	344 386	1.78 N 1.37 N	-11.7 +6.8	See Jan 11							
Apr 18 May 28 Jul 7*	40	0.40 N	-2.7	Herget 1977-EMP 80							
Jul 28 Aug 11 Sep 21	381 334	1.49 S 1.83 N	+4.0 -4.9								
Sep 24	11 148	1.67 N 1.17 N	-9.1 - 2.2	Early ITA-EMP 81							
Nov 15 Nov 17	375 * 375 *	0.54 N	+29.7 +30.0	Herget MPC 4372-EMP 81 Herget MPC 4372-EMP 81							
Dec 1 Dec 21 Dec 22	241	0.33 S 0.50 S	+5.9 +2.0	See Jan 2							

For the cases in the table not mentioned in the text above, the orbit differences are too small to decide by published recent observations. However, preliminary astrometry a few months beforehand could improve the predictions substantially for most of these events.

The occultations by Uranus were found by Arnold Klemola, Doug Mink, and Jim Elliot by scanning Lick Observatory plates. Their results for events for 1981 through 1984 were published in the *Astronomical Journal* early in 1981. General information about these events was published in O.N. 2 (10), 118. Elliot and other M.I.T. astronomers have found that occultations of even fainter stars than they've considered, often only 15th magnitude in V, can be recorded at major infrared observatories. Mink presented a long list of additional events in a talk at the American Astronomical Society's Division of Planetary Sciences in Pittsburgh, PA, in October.

A map showing my predicted paths of asteroidal occultations during 1982 in the U.S.A., southern Canada, and northern Mexico will be published in the 1982 January issue of *Sky and Telescope*. Mitsuru Sôma's world maps for early January occultations are included in this issue. Additional maps and some finder charts for the early 1982 events will be published in the next two issues of *O.N.*

Observing techniques and strategy have been discussed in considerable detail in articles in previous issues. Improved path predictions based on astrometry obtained a few to several days before possible North American events are usually available on recorded telephone messages at 312,259-2376 in Chicago, IL, and 501,771-0978 in Little Rock, AR. In the case when "last-minute" astrometry indicates an unusually favorable event, we will try to have a message broadcast on WWV. Since sky and Telescope is published more frequently than O.N., asteroidal occultation news and finder charts often will appear there first.

The generally poor coverage of the 1981 asteroidal occultations (so far) points out the need for more observers and better coordination; mobile observers are needed especially to fill in the gaps between fixed-site observers in widely separated cities. When you learn that last-minute astrometry shows that an asteroidal occultation is likely to occur in your area, pass the word on to other observers, especially those living in other cities and those with portable telescopes. The value of observations by two observers about a mile apart, to obtain independent confirmation of any observed events, and the need to practice locating the target star well before the occultation, can not be overemphasized!

Notes about Individual Events

Jan. 2: Sôma's world map was produced using an ephemeris based on Herget's orbit, but later texts mentioned above favored the more northern I.T.A. path plotted as a single solid line crossing the night part of the map.

Jan. 5: A special finder chart based mainly on Astrographic Catalog data is given to facilitate locating this faint uncatalogued star. The 8½-mag. SAO star nearby will help. But the gibbous moon only 2° away will make observation extremely difficult, perhaps possible only with large telescopes at major observatories. But so little is known about Chiron, the most distant asteroid, that some observational effort is encouraged.

Jan 11: Same note as for Jan. 2; the E.M.P. 81 path is the more likely.

Jan. 24: (532) Herculina probably has at least one large satellite, according to observations of the 1978 June occultation of a 6th-mag. star. Unfortunately, the small magnitude drop precludes visual observation.

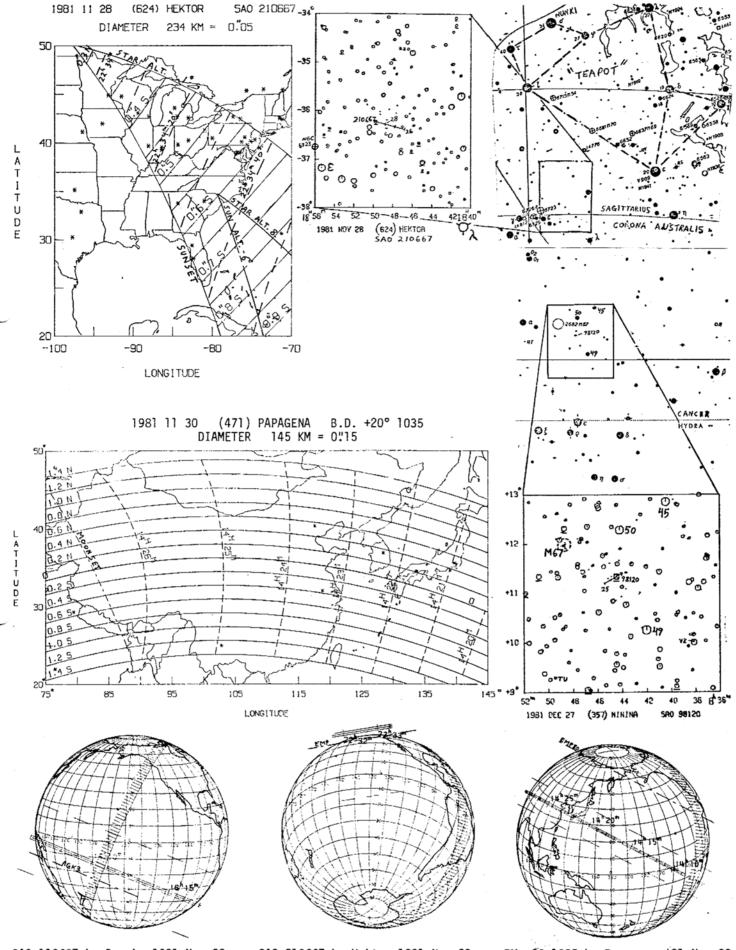
Notes about events later in 1982 will be included in future issues of O.N.

ERRATA

Ken Kelly reports the following errors and corrections in and to the Table of Ecliptic Variables, o.N. 2 (12), 162:

For SAO 97596, read 97496, For SAO 164829, read 164830.

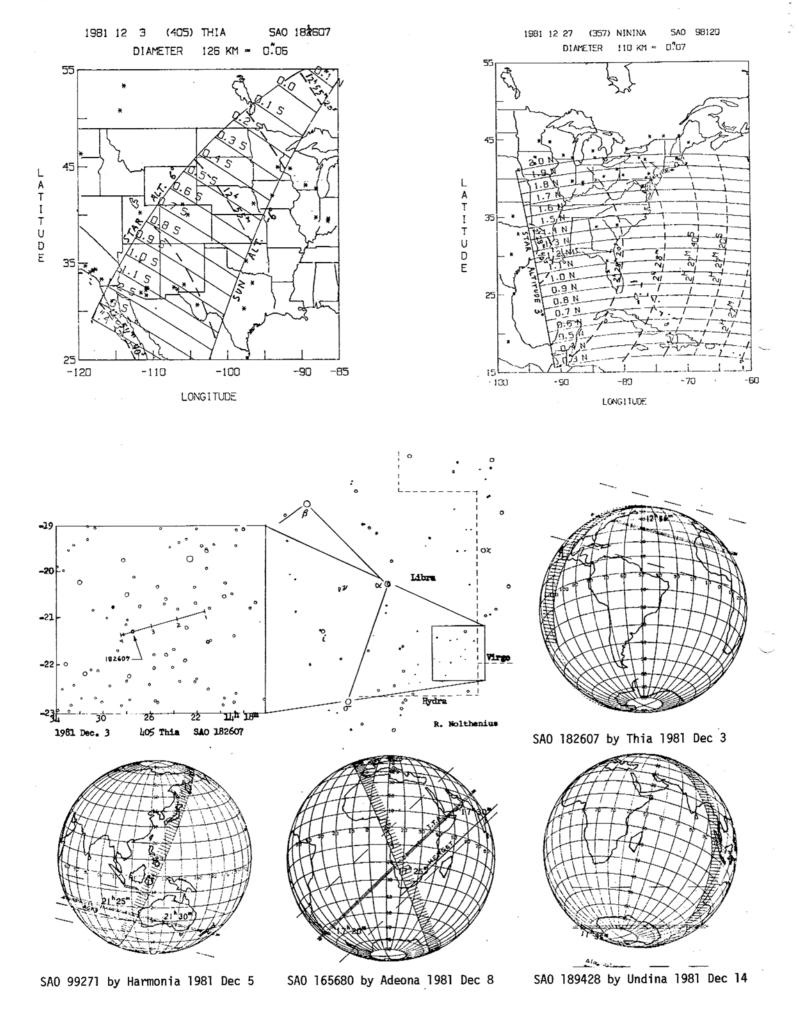
Dietmar Büttner has provided the value +28, to be substituted for the question mark in the entry for SAO 162511 (ZC 2825), in the article Erroneous Star Positions from Occultations, o.N. 2 (11), 137.

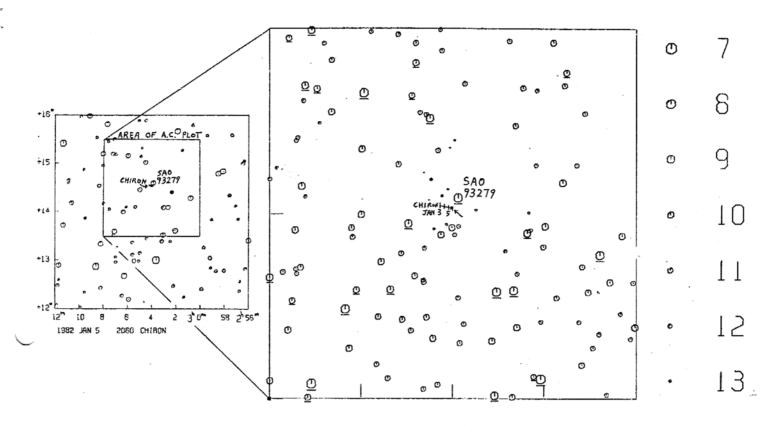


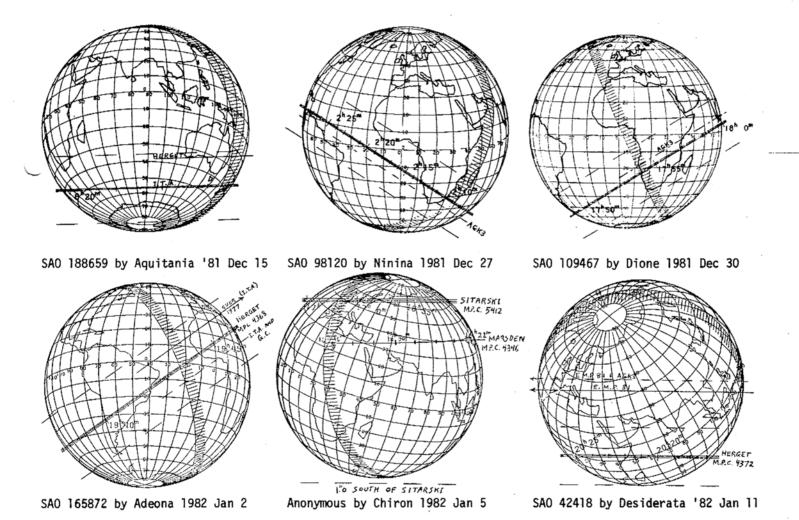
SAO 118607 by Psyche 1981 Nov 28

SAO 210667 by Hektor 1981 Nov 28

DM +20 1035 by Papagena '81 Nov 30







OUT OF FOCUS

Roger Giller

After a recent graze expedition, I had reason to criticise the way in which some supposedly mature and experienced observers had filled out a relative-

ly simple Station Report Form. This took the form of an article in our local society journal, *The Southern Observer*. In response, one of the team, Glen Dawes, of the Astronomical Society of New South Wales, decided that we needed a new Station Report Form. This was published in their journal, *Universe*. The form is reproduced below.

OOF GRAZING OCCULTATION OBSERVATION FORM

Sta Dat	N.B., form also used for observations of blue moons and aurora borealis (observed from Southern Hemisphere). State how many head of cattle were present. Date / / / / /(if known) Star No. / / / / / / (RS Catalogue No. preferred) Observer's Full Name / / / / Observing Conditions (A) / / Telescope Type / / / / / (Reflector, Refractor, Transit Instrument, etc.)													
	***TIMINGS													
U.T. (B) (E) (F) (C) (D) (G)														
(ERS		OBSERVATION										
HF			TION APPLIED? CERTAINTY METHOD DISCONTINUE											
	If further events	arı	re observed, additional forms can be obtained											
	EXHA	U S	STIVE O.O.F. CODES											
	(A) ODGERNANG CONDITIONS ON A CONTRACT OF THE													
	(A) OBSERVING CONDITIONS Yes/No If Yes, which of	۷)	Sure of event but possibly signal wrong star 3) Observer tripped over tripo	d										
	the following?	3)	Sure of event but observer is 4) Batteries failed in radio											
1)	Clear moonless day/night	-,	a compulsive liar 5) Batteries failed in recorde	r										
2)	Heavy fog	4)		Battery failed in car which										
3)	Raining		the star was sighted was running telescope											
4)	Snow / sleet	5)												
5)	Hail	۲.	was sighted 8) Police harassment (was bail											
6) 7)	Tornado	6)												
8)	Hurricane All of the above		active imagination 9) Milkman harassment											
0)	ATT OF CHE above		(D) TIMING METHOD 10) Neighbor complained / he couldn't stand the beeping	couldn't stand the beeping any										
	(B) EVENT CODE	1)		5										
1)	Disappearance	2)	Eye and Ear driveway	~										
2)	Reappearance	3)	Ear, nose, and throat 11) Neighbor's dog complained											
3)	Blink	4)	Tape recorder, voice, and - 12) Hand brake failed in car, o											
4)	Flash (Star rapid variable?)		a) Time signal rolled forward, running dow											
5)	Star failed to reappear		b) Water clock observer and his instrument											
6)	Star passed in front of moon		c) Hour glass 13) Observer run down by passin											
7)	(or seen through moon) A miss was seen / star veered		d) Egg timer car / did car swerve to avo e) Atomic clock missing you?	10										
′)	off at last minute	5)	Mickey Mouse wrist watch as- 14) Astronomer struck by lightr	ina										
8)	Moon disappeared / it set	٠,	sisted by time signal during freak storm	9										
	Total lunar eclipse		15) Swept away during flood cau	ised										
10)	Total solar eclipse		(G) IF OBSERVATION DISCONTIN- by same freak storm											
	Star went supernova		UED AT ANY STAGE: STATE 16) Earthquake (was it caused be	y										
12)	Moon went supernova	11	REASON conjunction of planets?)											
	(C) CERTAINTY	1)	Position of telescope was ad- 17) Observer kidnapped by passi iusted UFO	ng										
1)	Sure of event	2)	Radio adjusted due to loss of 18) Sun went supernova!!!											
		-,												

N.B., only reasons 17 and 18 will be accepted as an excuse for late lodging of results.

(E) PERSONAL EQUATION

$$Z = \sum \int \sqrt{z^2 + \log \left(\frac{\rho}{\delta}\right)} \cdot \int \sin \left(\epsilon \cos \left(\alpha\right)^{\sigma}\right)$$

Would all observers please supply own values for z, β , ρ , δ , κ , π , ϵ , α , ϕ (This request has been ignored in the past). For an explanation of the variables in this equation, refer to the *short Guide to OOF Graze Codes*, vI, 482 (At present permanently out of <u>feets</u> print).